

INITIAL UNSTEADY FREE CONVECTIVE FLOW PAST AN INFINITE VERTICAL PLATE WITH RADIATION AND MASS TRANSFER EFFECTS

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An exact solution and analysis of an initial unsteady two dimensional free convection flow, heat and mass transfer in the presence of thermal radiation along an infinite fixed vertical plate when the plate temperature is instantaneously raised, is presented. The fluid considered is a gray, absorbing emitting radiation but a non-scattering medium. Three cases have been discussed, in particular, namely, (i) when, the plate temperature is instantaneously raised to a higher constant value, (ii) when, the plate temperature varies linearly with time and (iii) when, the plate temperature varies non-linearly with time. A close form general solution for all the cases has been obtained in terms of repeated integrals of error functions. In two particular cases, the solutions in terms of the repeated integrals of error functions have been further simplified to forms containing only error functions. It is observed that for an increase in the radiation parameter N or a decrease in the Grashof number Gr or Gm, there is a fall in the velocity or temperature, but compared to the no radiation case or no diffusing species, there is a rise in the velocity and temperature of the fluid.

Key words: free convection flow, porous medium, heat transfer, mass transfer, MHD.

1. Introduction

The phenomenon of free convection arises in a fluid when temperature changes cause density variations leading to buoyancy forces acting on the fluid elements. This process of heat transfer is encountered in aeronautics, fluid fuel nuclear reactor and chemical engineering. Soundalgekar [1] presented an exact solution for the flow past an infinite vertical isothermal plate impulsively started in a viscous incompressible fluid and Patil and Soundalgekar [2] considered the same problem and presented an exact solution for an accelerated vertical infinite plate. Flow past a vertical oscillating plate with variable temperature has been discussed by Vighnesham *et al.* [3].

Korycki [4] stated that in textiles (i.e., industrial textiles, textiles designed for use under hermetic protective barrier, multilayer clothing materials, etc.) and in textile structures (i.e., needle heating in heavy industrial sewing), radiative heat transfer problems are encountered as well. Evidently, high temperature phenomena cannot be ignored. Therefore, it is more realistic to study the problem of an unsteady free convection flow with radiative heat transfer past a vertical porous plate. Chaudhary and Jain [5] studied the problem of an unsteady free convective flow past an impulsively started vertical surface, taking into account the effects of radiative heat transfer and suction. Heat and mass transfer effects on a moving plate in the presence of thermal radiation were studied by Muthukumarswamy *et al.* [6] using the Laplace technique. Chen [7] studied the problem of combined heat and mass transfer of an electrically conducting fluid in MHD natural convection, adjacent to a vertical surface with Ohmic heating. Chandran *et al.* [8] presented natural convection with ramped wall temperature. Several authors, namely, Muthucumaraswamy [9], Pathak *et al.* [10], Chandrakala [11] and Deka and Das [12], studied the problems of free convection flows, modeled

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under the assumptions of constant surface temperature, ramped wall temperature, or constant surface heat flux.

However, in practical problems when the plate temperature is raised, it may, initially increase linearly or non-linearly with time and therefore it is interesting to know how such initial thermal conditions of the plate affect the skin-friction, flow and heat transfer. Parihar [13] studied cases of such initial thermal conditions for a free convective flow in the absence of thermal radiation and species concentration. Furthermore, the free convection flows together with heat and mass transfer are of great importance in geophysics, aeronautics, and engineering. In several processes such as drying, evaporation of water at body surface, energy transfer in a wet cooling tower, and flow in a desert cooler, heat and mass transfer occur simultaneously. In view of such applications, several authors, namely, Chandrakala *et al.* [14], Muthucumaraswamy *et al.* [15], Muthucumaraswamy *et al.* [16], Prasad *et al.* [17], Narahari *et al.* [18], Rajesh [19], Brewster [20] and Abramowitz [21] investigated free convection flows with simultaneous heat and mass transfer phenomenon. However, the exact solutions to these problems have been obtained by few authors, namely Mebine *et al.* [22], Narahari *et al.* [23], Narahari *et al.* [18] and Rajesh [19].

In the present paper, an exact solution to the problem of the initial unsteady free convective flow and heat transfer along an infinite vertical plate, taking the thermal radiation and species concentration into account, has been obtained in three cases, namely, (i) when the plate temperature is instantaneously raised to a higher constant value, (ii) when the plate temperature varies linearly with time and (iii) when the plate temperature varies non-linearly with time.

The results are discussed in detail in section 5.

2. Formulation of the problem

Consider an initial unsteady two-dimensional free convective flow of viscous incompressible radiating fluid, with species concentration past an infinite vertical plate fixed in infinite mass of fluid which is otherwise at rest. At time $t \le 0$, the temperature of the plate and fluid is T_{∞} and for t > 0, the plate temperature is instantaneously raised to $T_w(t)$. It is assumed that the temperature difference between the plate and the fluid is small, so that the fluid properties may be taken as constant except for the influence of density variation in the body force term. The x-axis is taken along the plate in vertical upward direction and the y-axis is normal to it. Then, under the usual Boussinesq's approximations, the equations governing the flow, neglecting frictional heat, are given by

$$\frac{\partial u}{\partial t} = v \frac{\partial^2 u}{\partial y^2} + g\beta(T - T_{\infty}) + g\beta^*(C - C_{\infty}), \qquad (2.1)$$

$$\frac{\partial T}{\partial t} = \frac{\kappa}{\rho C p} \frac{\partial^2 T}{\partial y^2} - \frac{l}{\rho C p} \frac{\partial q_r}{\partial y}, \qquad (2.2)$$

$$\frac{\partial C}{\partial t} = D \frac{\partial^2 C}{\partial y^2},\tag{2.3}$$

with the following initial and boundary conditions

$$t < 0: \quad u = 0, \quad T = T_{\infty}, \quad C = C_{\infty} \quad \text{for all} \quad y,$$

$$t \ge 0: \quad u = 0, \quad T = T_{w}(t), \quad C = C_{w}(t) \quad \text{at} \quad y = 0,$$

$$u = 0, \quad T = T_{\infty}, \quad C = C_{\infty} \quad \text{as} \quad y \to \infty.$$
(2.4)

Here *u* is the velocity component in the *x* direction, *t* is the time, ν is the kinematic viscosity, *g* is the acceleration due to gravity, β is the coefficient of volume expansion. β^* is the volumetric coefficient of expansion with species concentration, *T* is the temperature of the fluid, T_{∞} is the temperature of the fluid far away from the plate. T_w is the plate temperature, κ is the thermal conductivity of the fluid, ρ is the density, C_p is the specific heat at constant pressure, q_r is the radiative heat flux in the y direction, *C* is the species concentration at infinity and *D* is the chemical molecular diffusivity.

The temperature of the plate and species concentration are now assumed to have the following form

$$T_w(t) = T_\infty \left(I + A t^{r-l} \right), \tag{2.5}$$

$$C_w(t) = C_\infty \left(l + A t^{r-l} \right) \tag{2.6}$$

where, $A = (u_0^2 / v)^{r-1}$, u_0 is a constant with dimension of velocity and $r \ge 1$ is an integer. It may be noted that r = I is the case of an isothermal plate (when the plate temperature is instantaneously raised to a constant higher temperature), r = 2 is the case of plate temperature varying linearly with time and $r \ge 3$ is the case of plate temperature varying non-linearly with time. These three particular cases have been discussed in particular.

We assume the Rosseland approximation [3], for the radiative heat flux, which leads to

$$q_r = -\frac{4\sigma}{3\kappa^*} \frac{\partial T^4}{\partial y}$$
(2.7)

where σ is the Stefan-Boltzmann constant and κ^* is the mean absorption coefficient.

If the temperature differences within the flow are sufficiently small such that T^4 may be expressed as a linear function of the temperature, then the Taylor series for T^4 about T_{∞} , after neglecting higher order terms, is given by

$$T^{4} = 4T_{\infty}^{3}T - 3T_{\infty}^{4}.$$
 (2.8)

In view of Eqs (2.7) and (2.8), Eq.(2.2) reduces to

$$\frac{\partial T}{\partial t} = \frac{\kappa}{\rho c_p} \frac{\partial^2 T}{\partial y^2} + \frac{16\sigma T_{\infty}^3}{3\rho c_p \kappa^*} \frac{\partial^2 T}{\partial y^2}.$$
(2.9)

Let us introduce the following non-dimensional quantities

$$y^* = \frac{yu_0}{v}, \quad u^* = \frac{u}{u_0}, \quad t^* = \frac{tu_0^2}{v}, \quad \theta = \frac{T - T_\infty}{T_w - T_\infty},$$
$$\Pr = \frac{\mu c_p}{\kappa}, \quad \operatorname{Gr} = \frac{vg\beta T_\infty}{u_0^3}, \quad N = \frac{\kappa^* k}{4\sigma T_\infty^3}, \quad \operatorname{Sc} = \frac{v}{D}$$
$$C^* = \frac{C - C_\infty}{C_w - C_\infty}, \quad \operatorname{Gm} = \frac{vg\beta^* C_\infty}{u_0^3}, \quad u_0 = \left(\frac{v}{t}\right)^{1/2}$$

where Pr is the Prandtl number, Gr the Grashof number, Gm the Grashof number for mass transfer, Sc the Schmidt number, N the radiation parameter and u_0 is a constant with dimension of velocity, in Eqs (2.1), (2.3), (2.4) and (2.9) which reduce to the following after dropping the astrisk sign

$$\frac{\partial u}{\partial t} = \frac{\partial^2 u}{\partial y^2} + \operatorname{Gr} t^{r-1} \theta + \operatorname{Gm} t^{r-1} C, \qquad (2.10)$$

$$\frac{\partial \theta}{\partial t} + \frac{\theta(r-I)}{t} = \frac{1}{\Pr} \left(\frac{3N+4}{3N} \right) \frac{\partial^2 \theta}{\partial y^2},$$
(2.11)

$$\frac{\partial C}{\partial t} + \frac{C(r-l)}{t} = \frac{l}{\operatorname{Sc}} \frac{\partial^2 C}{\partial y^2},$$
(2.12)

with the following initial and boundary conditions

$$t < 0: u = 0, \quad \theta = 0, \quad C = 0 \text{ for all } y,$$

 $t \ge 0: u = 0, \quad \theta = 1, \quad C = 1 \quad \text{at } y = 0,$
 $u = 0, \quad \theta = 0, \quad C = 0 \quad \text{as } y \to \infty.$
(2.13)

The partial differential Eqs (2.10), (2.11) and (2.12) with initial and boundary conditions (2.13) can be solved either by the Laplace–transform technique or by using similarity transformation.

3. Solution of the problem

We choose the latter technique and introduce the following similarity transformations

$$u(y,t) = t^{P} f(\eta),$$

$$\theta(y,t) = \phi(\eta),$$

$$C(y,t) = \psi(\eta),$$
and
$$\eta = \frac{y}{2\sqrt{t}},$$
(3.1)

into Eqs (2.10), (2.11) and (2.12) which reduce to the following ordinary differential equations

$$f'' + 2\eta f' - 4rf + 4Gr.\phi + Gm.\psi = 0, \qquad (3.2)$$

$$\frac{1}{\Pr.\alpha}\phi'' + 2\eta\phi' - 4(r-I)\phi = 0, \qquad (3.3)$$

$$\frac{1}{Sc}\psi'' + 2\eta\psi' - 4(r-I)\psi = 0, \qquad (3.4)$$

with the following boundary conditions

$$\eta = 0; \quad f = 0, \quad \phi = 1, \quad \psi = 1,$$

$$\eta \to \infty; \quad f = 0, \quad \phi = 0, \quad \psi = 0.$$
 (3.5)

Here, $\alpha = \frac{3N}{3N+4}$, and a prime denotes differentiation with respect to η .

The solutions of Eqs (3.2), (3.3) and (3.4) can be easily obtained in terms of repeated integrals of complementary error functions (see, e.g. Abramovitz and Stegun [21]). Equation (3.3) and (3.4), being independent of (3.2), is first solved, satisfying the boundary conditions (3.5). Then a particular integral of Eq.(3.2), in view of the solution of Eqs (3.3) and (3.4), is obtained by the method of undetermined coefficients (Bansal [24]) and finally, the complete solution of Eq.(3.2), satisfying the boundary conditions (3.5), is obtained. These solutions are as follows

$$\theta = 4^{r-1} \Gamma r \ i^{(2r-2)} \mathrm{erf}_{\mathrm{c}}(\eta \sqrt{\mathrm{Pr}.\alpha}), \tag{3.6}$$

$$C = 4^{r-1} \Gamma r \ i^{(2r-2)} \operatorname{erf}_{c}(\eta \sqrt{\operatorname{Sc}}), \tag{3.7}$$

$$\mathcal{U} = \frac{t^{r} 4^{r} \Gamma r \operatorname{Gr}}{(l - \operatorname{Pr}.\alpha)} \Big[i^{2r} \operatorname{erf}_{c}(\eta \sqrt{\operatorname{Pr}.\alpha}) - i^{2r} \operatorname{erf}_{c}(\eta) \Big] + \frac{t^{r} 4^{r} \Gamma r \operatorname{Gm}}{(l - \operatorname{Sc})} \Big[i^{2r} \operatorname{erf}_{c}(\eta \sqrt{\operatorname{Sc}}) - i^{2r} \operatorname{erf}_{c}(\eta) \Big].$$
(3.8)

Here, the function $i^n \operatorname{erf}_{c}(z)$ is the repeated integral of the complementary error function and is given by

$$i^{n} \operatorname{erf}_{c}(z) = \frac{2}{\sqrt{\pi}} \int_{z}^{\infty} \frac{(t-z)^{n}}{n!} e^{-t^{2}} dt, \qquad n = 0, 1, 2, \dots$$
(3.9)

The solution (3.8) is valid for $Pr.\alpha \neq l$ and $Sc \neq l$. In the case of $Pr.\alpha = l$ or Sc=l, on taking the limiting values, one can easily obtain the following expressions for the velocity.

Case (i): When $Pr \cdot \alpha = l$ and $Sc \neq l$

$$\mathcal{U} = t^{r} 4^{r-l} \Gamma r \operatorname{Gr} \left[i^{2r-2} \operatorname{erf}_{c}(\eta) - 4r i^{2r} \operatorname{erf}_{c}(\eta) \right] + \frac{t^{r} 4^{r} \Gamma r \operatorname{Gm}}{(l-\operatorname{Sc})} \left[i^{2r} \operatorname{erf}_{c}(\eta \sqrt{\operatorname{Sc}}) - i^{2r} \operatorname{erf}_{c}(\eta) \right].$$
(3.10)

Case (ii): When $Pr.\alpha \neq l$ and Sc=l

$$\mathcal{U} = \frac{t^{r} 4^{r} \Gamma r \operatorname{Gr}}{(1 - \operatorname{Pr}.\alpha)} \Big[i^{2r} \operatorname{erf}_{c}(\eta \sqrt{\operatorname{Pr}.\alpha}) - i^{2r} \operatorname{erf}_{c}(\eta) \Big] + t^{r} 4^{r-1} \Gamma r \operatorname{Gm} \Big[i^{2r-2} \operatorname{erf}_{c}(\eta) - 4r \cdot i^{2r} \operatorname{erf}_{c}(\eta) \Big].$$
(3.11)

Case (iii): When $Pr.\alpha = l$ and Sc=l

$$\mathcal{U} = t^r 4^{r-1} \Gamma r \left(\text{Gr} + \text{Gm} \right) \left[i^{2r-2} \text{erf}_{c}(\eta) - 4r i^{2r} \text{erf}_{c}(\eta) \right].$$
(3.12)

It may be noted that Gm = 0 is the case of absence of species concentration and $N \rightarrow \infty$ or $\alpha \rightarrow l$ is the case of absence of the radiative heat flux. In the case of absence of both the species concentration and radiative heat flux, the above expressions for velocity reduce to those obtained by Tak and Parihar [12].

4. Skin-friction coefficient

The non-dimensional skin-friction at the plate is determined by the formula.

$$C_{f} = \frac{\mu \left(\frac{\partial u}{\partial y}\right)_{y=0}}{\frac{l}{2}\rho u_{0}^{2}} = 2t^{r-\frac{l}{2}} \frac{\Gamma r}{\Gamma(r+1/2)} \left[\frac{\mathrm{Gr}}{(l+\sqrt{\mathrm{Pr}\,\alpha})} + \frac{\mathrm{Gm}}{(l+\sqrt{\mathrm{Sc}\,})}\right].$$
(4.1)

In the case of no radiation ($\alpha = I$) and no mass transfer (Gm = 0), the expression for C_f becomes

$$C_f = 2t^{r-\frac{1}{2}} \frac{\Gamma r}{\Gamma(r+1/2)} \left[\frac{\mathrm{Gr}}{(1+\sqrt{\mathrm{Pr}\,\alpha})} \right]. \tag{4.2}$$

5. Results and discussion

It may be noted from the non-dimensional Eqs (3.2) to (3.5) that the present problem is described by six parameters, namely Gr, Gm, Pr, Sc, r and α (or N). Since the effects of variation of Pr and Sc on the flow are well known, all numerical calculations have been performed for Pr = 0.71 and Sc = 0.6. It may be noted that α =1 (or N $\rightarrow \infty$) is the case when the thermal radiation is neglected and Gm =0 is the case of absence of species concentration. The value of the parameter r determines whether the temperature and concentration at the plate are constant or vary with time (see, Eqs (2.5) and (2.6)). We have chosen r=1, 2 and 3 which describes three different physical situations of the flow and discussed the results as follows:

Case (i) *r*=1 (Isothermal and isosolutal plate):

This corresponds to the situation of an initial unsteady free convective flow when the temperature and concentration of the plate are raised to attain higher constant values instantaneously. In this case, we have drawn three-dimensional graphs for velocity distribution u(y, t) for different values of Gr, Gm and N, in Figs 1-4. Figure 1 depicts the developing velocity profiles in the presence of thermal radiation and diffusing species when Gr=4, Gm=4 and N=3, whereas Fig.2 depicts the case when thermal radiation is neglected (α =1 or N $\rightarrow \infty$), Fig.3 depicts the case when species concentration is absent (Gm=0) and Fig.4, when both radiation and concentration are absent (α =1 and Gm=0).



Fig.1. Variation of velocity with y and t in the case of isothermal plate (r=1).



Fig.2. Variation of velocity with y and t in the case of isothermal plate (r=1) in the absence of radiation.



Fig.3. Variation of velocity with y and t in the case Fig.4. Variation of velocity with y and t in the case of isothermal plate (r=1) in the absence of mass transfer.



of isothermal plate (r=1) in the absence of both mass transfer and radiation.

A comparison of Fig.1 with Figs 2, 3 and 4 reveals that there is a rise in velocity due to the presence of radiation and/or diffusing species as compared to the cases of no radiation and/or no diffusing species.

In Figs 7 and 8, two dimensional graphs have been drawn to depict the variation of velocity u with similarity variable η for different values of the radiation parameter N, in the presence and absence of mass transfer, respectively, at t=0.4. It may be observed from these figures that for an increase in parameter N, there is a fall in the velocity whereas for an increase in Gr or Gm, there is a rise in the velocity. Further, the absence of radiation heat transfer and/or mass transfer causes the velocity to decrease which is in conformity with the observation drawn from 3D Figs 1-4.

Case (ii) *r*=2 (Plate temperature and concentration vary linearly with time):

This is the case of an initial unsteady flow when the temperature and concentration at the plate are both linear functions of time. In this case, Fig.5 displays a 3D graph for velocity distribution u(y, t) for Gr=4, Gm=4, N=3 and Fig.9 displays a 2D graph for velocity u against η . Comparing Figs 1 and 5, it may be observed that there is a fall in the velocity when the plate temperature and concentration vary linearly with time for a fixed time t < 1. The effects of other parameters are same as those observed in case (i).

Case (iii) *r*=3 (Plate temperature and concentration vary non-linearly with time):

This is the case of an initial unsteady flow when the temperature and concentration at the plate are both quadratic functions of time. In this case, Fig.6 displays a 3D graph for velocity distribution u(y, t) for Gr=4, Gm=4, N=3 and Fig.10 displays a 2D graph for velocity u against η . Comparing Fig.6 with Figs 1and 5, it may be observed that there is a further fall in the velocity when the plate temperature and concentration vary non-linearly with time for a fixed time t < I, as compared to cases (i) and (ii).



Fig.5. Variation of velocity with y and t in the case of plate temperature varying linearly with time (r=2).



Fig.7. Variation of velocity (*u*) with similarity variable (η) in the case of isothermal plate (*r*=1).



Fig.6. Variation of velocity with y and t in the case of plate temperature varying non-linearly with time (r=3).



Fig.8. Velocity variation of velocity (*u*) with similarity variable (η) in the case of isothermal plate (r=1) in the absence of mass transfer.



variable (η) in the case of plate temperature varying linearly with time (r=2).

Fig.9. Variation of velocity (u) with similarity Fig.10. Variation of velocity (u) with similarity variable (η) in the case of plate temperature varying non-linearly with time (r=3)

1.50

1.00

η

2.00

0.50

Finally, the temperature function θ and concentration function C have been plotted against the similarity variable η for different values of r and N in Figs 11 and 12, respectively.

It is noted from Fig.11 that for all values of r there is a rise in temperature due to the presence of thermal radiation as compared to the case of no radiation. Further, the temperature decreases with increasing values of N. As $N \rightarrow \infty$ (no radiation case), the limiting temperature profile is also shown in Fig.11, which is the minimum temperature profile as compared to the cases of radiation (N = 3, 5, 15). We also observe that the temperature decreases as r increases.

In Fig.12, species concentration profiles are shown against η , for various values of r. We observe that the species concentration decreases as r increases.



Fig.11. Temperature profile for Pr=0.71and Fig.12. Species concentration profile for Sc=0.6 and various values of N. various values of r.

N=3

N-5

- N-15

 $\alpha = 1$

T :6 6

II :4 4

2.50

3.00

Gm Gr

We have also calculated numerical values of the skin-friction C_f that are listed in Tabs 1-3. We observe from Tab.1 that a decrease in the radiation parameter N or an increase in the mass transfer parameter Gm or Grashof numbers Gr or time t, leads to an increase in the skin-friction. We also observe from Tabs 1, 2 and 3 that the skin-friction decreases as r increases.

In Tab.4, numerical values of the skin-friction C_f , in the absence of radiation, are listed. We compare these values of C_f with those in Tabs 1-3 and conclude that the value of the skin-friction in the presence of radiation is greater than that in the non-radiation case, other parameters being fixed.

Hence, the presence of radiation causes a rise in the skin-friction when other parameters are fixed. Also, we observe that the presence of species concentration causes a rise in the skin-friction when other parameters are fixed.

In Tab.3, numerical values of the skin-friction C_f , in the absence of radiation, are listed. We compare these values of C_f with those in Tabs 1-2 and conclude that the value of skin-friction is less than that in the non-radiation case. Hence, the presence of radiation causes a fall in the skin-friction at various values of t and Gr. Also observe that the presence of species concentration causes a rise in the skin-friction at various values of t and Gr.

Table 1. Numerical values of skin-friction of isothermal plates for Pr=0.71, Sc=0.6.

Fable 2. Numerical	values of	skin-fric	tion in	the ca	se of
plate temp	varying li	nearly for	r Pr=0.	71, Sc	=0.6.

Gr	Gm	Ν	Т	C.F.
2	2	3	0.2	2.324035
4	4	3	0.2	4.64807
6	6	3	0.2	6.972105
2	2	5	0.2	2.291746
4	4	5	0.2	4.583491
6	6	5	0.2	6.875237
2	2	15	0.2	2.25419
4	4	15	0.2	4.50838
6	6	15	0.2	6.76257
2	2	3	0.4	3.286682
4	4	3	0.4	6.573363
6	6	3	0.4	9.860045
2	2	5	0.4	3.241018
4	4	5	0.4	6.482036
6	6	5	0.4	9.723053
2	2	15	0.4	3.187906
4	4	15	0.4	6.375812
6	6	15	0.4	9.563718
2	0	3	0.4	1.678092
4	0	3	0.4	3.356183
6	0	3	0.4	5.034276
2	0	5	0.4	1.632428
4	0	5	0.4	3.264855
6	0	5	0.4	4.897283
2	0	15	0.4	1.579316
4	0	15	0.4	3.158632
6	0	15	0.4	4.737948

Gr	Gm	Ν	Т	C.F.
2	2	3	0.2	0.3098713
4	4	3	0.2	0.6197426
6	6	3	0.2	0.929614
2	2	5	0.2	0.3055661
4	4	5	0.2	0.6111322
6	6	5	0.2	0.9166982
2	2	15	0.2	0.3005587
4	4	15	0.2	0.6011174
6	6	15	0.2	0.901676
2	2	3	0.4	0.8764485
4	4	3	0.4	1.752897
6	6	3	0.4	2.629345
2	2	5	0.4	0.8642714
4	4	5	0.4	1.728543
6	6	5	0.4	2.592814
2	2	15	0.4	0.8501083
4	4	15	0.4	1.700217
6	6	15	0.4	2.550325
2	0	3	0.4	0.4474911
4	0	3	0.4	0.8949822
6	0	3	0.4	1.342474
2	0	5	0.4	0.4353141
4	0	5	0.4	0.8706281
6	0	5	0.4	1.305942
2	0	15	0.4	0.421151
4	0	15	0.4	0.8423019
6	0	15	0.4	1.263453

r	Gr	Gm	C.F.
1	2	2	3.157801
1	4	4	6.315601
1	6	6	9.473401
1	2	0	1.549211
1	4	0	3.098421
1	6	0	4.647632
2	2	2	0.8420802
2	4	4	1.68416
2	6	6	2.526241
2	2	0	0.4131228
2	4	0	0.8262456
2	6	0	1.239368

Table 3. Numerical values of skin-friction for Pr=0.71, Sc=0.6, t=0.4 and $\alpha=1$ (no radiation case).

Table 4. Numerical values of skin-friction for Pr=0.71, Sc=0.6, t=0.4 and $\alpha=1$ (no radiation case).

r	Gr	Gm	C.F.
1	2	2	3.157801
1	4	4	6.315601
1	6	6	9.473401
1	2	0	1.549211
1	4	0	3.098421
1	6	0	4.647632
2	2	2	0.8420802
2	4	4	1.68416
2	6	6	2.526241
2	2	0	0.4131228
2	4	0	0.8262456
2	6	0	1.239368
3	2	2	0.2694657
3	4	4	0.5389313
3	6	6	0.8083969
3	2	0	0.1321993
3	4	0	0.2643986
3	6	0	0.3965979

Nomenclature

- C species concentration
- C_f skin friction coefficient
- C_p specific heat at constant pressure
- C_{∞} concentration at infinity
- D chemical molecular diffusivity

- Gm Grashof number for mass transfer
- Gr Grashof number
- g acceleration due to gravity
- $i^{n} \operatorname{erf}_{c}(z)$ repeated integral of complementary error function
 - N radiation parameter
 - Pr Prandtl number
 - $Sc \ -Schmidt \ number$
 - T temperature of the fluid
 - T_{∞} temperature of the fluid far away from the plate
 - T_w plate temperature
 - t time
 - u velocity component
 - u_0 constant with dimension of velocity
 - q_r radiative heat flux
 - v kinematic viscosity
 - β coefficient of volume expansion
 - β^* volumetric coefficient of expansion with species concentration
 - θ dimensionless temperature
 - κ thermal conductivity of the fluid
 - κ^* mean absorption coefficient
 - ρ density
 - σ Stefan-Boltzmann constant

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